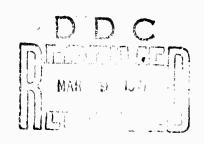
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SCIENTIFIC REPORT (Report No. 21)

FREQUENCY-DEPFNDENT AMPLITUDE-DISTANCE CURVE FOR P-WAVES FROM 87° TO 110°

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#### Summary

This work is an attempt to clarify the nature of the amplitudedistance curve for P-waves between 87° and 110°, using the spectral amplitudes of earthquakes in the Indonesian region recorded at the Swedish and Finnish seismograph stations. At the present time the results are inconclusive, because even after allowing for station and source terms there is a large unexpected scatter.

## Introduction

beyond 95° and which have passed through the deeper part of the mantle has been observed for a long time. As the quality and distribution of seismographs have increased, it has been possible to make more detailed studies. Gutenberg (1960) with limited data demonstrated the frequency dependence of the amplitude decrease, while Sacks (1966) with better data illustrated the effect very clearly. These results, however, are more qualitative than quantitative and cannot be used to test hypotheses on the nature of the core-mantle boundary region which produces the shadowing effect. Also these studies use a few earthquakes and do not consider the effects of station geology on the results. More recently, Alexander and Phinney (1966) have worked with long-period waves in the shadow region, but their data has large scatter, they do not consider station effects and they do not combine data from different earthquakes.

Recently, Carpenter, Marshall and Douglas (1957) and Cleary (1967) have worked on the amplitude-distance curve between 30° and 102° and have

used a joint analysis method described by Carpenter et al. This method allows the combination of different earthquakes, finds corrections for the station effect and produces an amplitude-distance curve independent of earthquakes and stations. These amplitude-distance curves are valid for short-period vertical-component records of about 1 sec period, but there is some disagreement between them at distances beyond 90° probably because of the different methods of measuring amplitudes. Both authors have little data over 95° and neither investigate the effect of frequency on the amplitude-distance curve. The present work has been done to attempt to clarify the frequency dependence of the amplitude-distance curve using the joint analysis technique to combine data from many earthquakes and many stations.

The observational material used in the present investigation consists of short-period vertical-component records of P-waves from the network of Swedish and Finnish stations for a number of earthquakes in the Indonesian archipelago.

#### Analytical method

At teleseismic distances we can express the frequency-dependent amplitude A of the body waves in the form

$$A = B.R.S \tag{1}$$

where B is the frequency-dependent source function which includes the effect of the crust and upper mantle at the source, R is the transmission coefficient for passage through the mantle which includes the effects of reflection and diffraction by the core-mantle boundary region (if the wave concerned is affected by this region), the effect of transmission at any boundaries and the effect of geometrical spreading of the waves, and

finally S is the receiver function which includes the effects of the station seismographic response curve and the crust and the upper mantle below the station. Each of B, R and S also includes the effect of the anelastic dissipation and scattering by inhomogeneities in the regions concerned. B and S vary with azimuth and also with angle of incidence on the surface. To the extent that the lower mantle is inhomogeneous R is dependent on the particular path through the mantle.

In our problem then, we have selected the stations and earthquakes such that we make the assumption that B and S vary little over the small variation of azimuthal angles and small variation of angles of incidence involved (this assumption may not be valid!). R will apply to the mantle between Indonesia and Fennoscandia and to the core-mantle boundary region under Central Asia. All of A, B, R and S are frequency-dependent and complex.

If we use the base ten logarithms of these quantities then we have

$$a = b + r + s \tag{2a}$$

where a =  $log_{10}|A|$ , b =  $log_{10}|B|$  etc. (|A| is the amplitude of the complex A) and

For any particular measurement of a we have

$$a = b + r + s + \varepsilon \tag{3}$$

where  $\varepsilon$  is an error term which includes the inaccuracies of measurement of a and the effect of inadequacies of the model we have set up. This formulation is the same as derived by Carpenter et al. (1967) except that our a is the log of the spectral amplitude and not  $\log_{10} \left(\frac{A}{T}\right)$  and in our model the azimuths and angles of incidence are very limited in range.

Following Carpenter et al. we find that if we make a number of observations of a at a number of stations for a number of earthquakes we can obtain estimates of b, r and s. If we denote by subscript i the particular earthquake considered then b; is the source term of the i<sup>th</sup> earthquake. Similarly, if we denote by subscript j the particular station considered then s; is the station (crustal + seismograph) function for the i<sup>th</sup> station. Finally, if we divide the distance range into intervals over which the amplitude-distance curve is assumed constant and we denote by k the k<sup>th</sup> such interval, then r<sub>k</sub> is the mantle transfer function for this distance range. If some estimate r<sub>e</sub> of the amplitude-distance curve is available, then we can subtract r<sub>e</sub> from both sides of equation (3) and r<sub>k</sub> is considered as the actual difference of r and r<sub>e</sub>. When r<sub>e</sub> is a reasonable approximation to r, the constancy of r<sub>k</sub> over a distance interval is a less imposing condition and yet we retain the flexibility of the histogram representation.

Thus if earthquake i is observed at station j and the separation of the two is in distance range k, the observed amplitude a ijk may be expressed in the form

$$a_{ijk} = b_i + r_k + s_j + \epsilon_{ijk}$$
 (4)

For Nre observations of  $a_{ijk}$  from Nep epicentres at some or all of Nst stations using Npa distance ranges, we have a set of Nre linear equations for  $a_{ijk}$ . We have Nep unknowns  $b_i$ , Npa unknowns  $r_k$ , Nst unknowns  $s_j$  and Nre unknowns  $s_{ijk}$ . If we remove from  $b_i$ ,  $r_k$  and  $s_j$  their respective averages so that

$$C = \overline{b}_{i} + \overline{r}_{k} + \overline{s}_{j}$$

$$b_{i} - \overline{b}_{i} = b_{i}'$$

$$r_{k} - \overline{r}_{k} = r_{k}'$$

$$s_{j} - \overline{s}_{j} = s_{j}'$$
(5)

then the new  $b_i$ ',  $r_k$ ' and  $s_j$ ' averaged over i, k and j respectively are zero. Hence we have Nre + 3 equations:

Nre equations

$$a_{ijk} = C + b_{i}' + r_{k}' + s_{j}' + \epsilon_{ijk}$$
and 3 equations
$$\sum_{i} b_{i}' = 0, \quad \sum_{k} r_{k}' = 0 \quad \text{and} \quad \sum_{j} s_{j}' = 0$$

$$(6)$$

We can henceforth drop the primes on bi, rk and si.

We can represent these equations in the matrix formulation

$$P = QX + E \tag{7}$$

where P is the row vector of  $\mathbf{a_{ijk}}$  in some order, E is the error row vector of  $\mathbf{c_{ijk}}$  in the same order as  $\mathbf{a_{ijk}}$ , X is the column vector (C,b<sub>1</sub>,b<sub>2</sub>...b<sub>Nep</sub>, r<sub>1</sub>, r<sub>2</sub>...r<sub>Npa</sub>, s<sub>1</sub>, s<sub>2</sub>...s<sub>Nst</sub>) and Q is the matrix of indicator variables such that if the n<sup>th</sup> element of P is  $\mathbf{a_{ijk}}$  then the n<sup>th</sup> row of Q multiplied by X gives C + b<sub>i</sub> + r<sub>k</sub> + s<sub>j</sub> and the last three rows of Q when multiplied by X give equations (6).

It is possible to solve this matrix equation by the least squares method to minimise |E| and get an estimate of X and hence of C and of b,  $r_k$  and  $s_j$ .

The least squares estimate for X is given by

$$x = (Q^{T}Q)^{-1}(Q^{T}P)$$
 (8)

where  $Q^T$  is the transpose of Q and  $(Q^TQ)^{-1}$  is the inverse of  $Q^TQ$ , i.e.  $(Q^TQ)^{-1}(Q^TQ) = I$ , the identity matrix. Problems may arise with calculation of  $(Q^TQ)^{-1}(Q^TP)$  and these problems are discussed by Anderssen (1969). In the present work straightforward matrix inversion

was used to form  $(Q^TQ)^{-1}$  and the difference  $(Q^TQ)^{-1}(Q^TQ)$  -I was used as a guide to the accuracy of the inversion of  $Q^TQ$ . Since Q is composed of integer indicator variables,  $Q^TQ$  can be calculated exactly and is not affected by computational rounding errors. If there are a sufficient number Ni of linearly independent equations (4), then a solution X of equation (7) can be found. Ni + 3 must be greater than 1 + Nep + Npa + Nst, and the greater Ni the better the statistical estimate of X.

## Observational material

The stations used are those of the high quality Swedish and Finnish networks situated on the relatively homogeneous Fennoscandian shield (table 1). The earthquakes used occurred in the Indonesian area between the beginning of 1963 and the end of 1968 (table 2). See also figure 1. The particular earthquakes selected were such that the signal-to-noise ratio was generally good, the amplitude of the signal was sufficient to make further analysis worthwhile and the energy of the signal was concentrated near the onset. Any selection of the data will affect the final result as the criteria used are subjective. If, for example, a record is rejected because of low signal-to-noise ratio - then it may be that the noise level is high or that the amplitude level is low. However, some selection must be made and the criteria used seem reasonable.

The stations and earthquakes are related so that for any one earthquake the stations cover an azimuthal range of less than 10° and that for one station the earthquakes cover a back azimuthal range of less than about 20° (cf figure 2). For the range 90°-110° epicentral distance, the angle of approach of the seismic P-wave changes little. So for each earthquake the station net covers a small solid angle and for each station the

earthquake epicentres cover a small solid angle. As we shall see later, these conditions should make the joint analysis method suitable for analysing the data.

The eleven Swedish and Finnish stations originally chosen are given in table 1. Of these SOD was later rejected, because of the nonstability of its amplification curve, and UDD, which because of its later construction recorded only four of the earthquakes (two on the earlier Grenet instrument and two on the later installed Benioff).

Sixteen earthquakes were initially selected, listed in table 2.

Eleven of these earthquakes lie in a narrow back azimuthal range from

Scandinavia and the other five are outside this band. The latter five are

treated as suspect, as the station terms may vary too much with large

changes in back azimuth. Of the original 176 possible records, 109 were

selected and digitised. 14 records were not available, 20 were at too

short epicentral distances and 33 were rejected because the signal-to
noise ratio was too low or the record amplitude was not large enough to

make Fourier analysis worthwhile. Figure 3 shows a typical record.

For each earthquake the epicentral distance, azimuth and back azimuth to the stations of the net were calculated. The epicentral distances were corrected for depth of focus using the results of Buchbinder (1968). These corrections are such that all the earthquakes can be considered as surface focus events with regard to the amplitude-distance curve.

#### Experimental method

For the records from each earthquake a suitable record length was chosen, either 20, 30 or 40 sec, and this length was such that the main

part of the energy was in the earlier portion of the record and at the end of the record the amplitude was much smaller or reduced to near noise level. This selection of record lengths should minimise the effects of truncation of the record. The start of the record was taken just before the apparent onset of the arriving P-wave.

The records were photographically enlarged four or five times. Then the top and bottom of the trace were digitised on a DMac pen follower and the data were converted to cards. They were then interpolated to the desired interpolation interval:  $\frac{20}{256}$  sec for 20 sec records,  $\frac{30}{512}$  sec for 30 sec records and  $\frac{40}{512}$  sec for 40 sec records. The average of the two traces was taken and the Fourier transform of the average computed in the form of amplitude and phase spectra. The theory of the spectral analysis of digitised seismic data is well covered by Huang (1966). If the seismic trace is the time function f(t), then the computed Fourier spectrum is given by  $F(v_n)$  where

$$F(v_n) = \frac{1}{m} \sum_{k=0}^{m-1} f(k\Delta t) e^{-\frac{2\pi i n k \Delta t}{T}}$$
(9)

T is the length of the record,  $v_n$  is the n<sup>th</sup> frequency in cycles per sec and  $v_n = \frac{n}{T}$  where n runs from 0 to m, m is the number of digitised points, at is the digitising interval and mat = T, and finally for most efficient computation m is a power of 2 (in our case m = 256 or 512). We avoid aliasing by using a digitising interval sufficiently small so that the amplitude spectra are neglible for frequencies above the folding frequency  $\frac{1}{2\Lambda t}$ .

Various methods of windowing were considered but none was applied

as none seemed suitable. Using longer record lengths with low cut-off amplitudes should minimise the effect of truncating the records.

One record of average quality was photographically enlarged and digitised separately three times. The agreement between the three amplitude spectra is very good as shown in figure 4. The maximum variation throughout most of the frequency range was 0.25 units compared with the maximum amplitude of 4 units. For the larger amplitude components the difference is less than 8%. For better quality records the agreement should be better and the opposite for poorer quality records.

The amplitude spectra have been smoothed using a 3-point smoothing with  $\frac{1}{4}$ ,  $\frac{1}{2}$ ,  $\frac{1}{4}$  weighting for the 20 sec records, a 3-point smoothing with  $\frac{1}{3}$ ,  $\frac{1}{3}$ ,  $\frac{1}{3}$  weighting for the 30 sec records and a 5-point smoothing with  $\frac{1}{8}$ ,  $\frac{1}{4}$ ,  $\frac{1}{4}$ ,  $\frac{1}{4}$ ,  $\frac{1}{8}$  weighting for 40 sec records. This smoothing should improve the consistency of the results and also make the comparison of the spectra from different length records more meaningful. This smoothing is not windowing but an attempt to smooth the insignificant fluctuations in the amplitude spectra.

All the spectra obtained are divided by the instrument magnification factor at 1 sec, and so the amplification curves are normalised at 1 sec. Also the inverse of the magnification factors for the photo enlargement is applied such that the amplitude is in units of 0.1 microns and after we have taken the  $\log_{10}$  of the amplitude spectra, we add one to the results, i.e. the log (amplitude) of the spectra is such that the amplitude is measured in 0.01 microns. PcP is always included in the pulse, and if the earthquake is shallow, pP is included in the record. If

the earthquake is deeper, pP either does not affect the record or is small and appears near the end of the record.

## Computations and results

In the experimental work we find estimates of a ijk for various earthquakes and stations. Then we apply the analytical method of joint analysis to estimate the amplitude-distance curve and the station terms. A computer program to solve equation (7) and find X in the form of equation (8) has been developed. The program calculates the station and source terms s<sub>j</sub> and b<sub>i</sub> and also the amplitude-distance curve using a histogram of 2° intervals. See Appendix.

The results are very poor. In figure 5 we show a plot of the raw data from which is subtracted the appropriate source and station terms and the constant introduced in equation (5). Even though the station and source terms are allowed for, the scatter is very high and certainly no amplitude decrease is seen beyond 90° - as would be expected. This behaviour seems to come from the data and not from the inversion program. So far no explanation has been found for the anomalous behaviour of the results. The data has been divided into smaller groups of earthquakes with narrower azimuthal and distance ranges but there is no substantial improvement in the results. It is possible that the model we have set up is based on invalid assumptions on the nature of the source and station functions.

As an example we present the amplitude data for 1 sec period for all stations which recorded the earthquake on the 29 July, 1968. We list the stations in order of azimuth (epicenter to station) and epicentral distance. It is obvious that no clear pattern emerges.

Station	Azimuth	log <sub>10</sub> (Amp) at 1 sec		
KLS	329°1	- 0.1664		
GOT	331.5	- 0.1292		
NUR	331.7	- 0.1053		
UPP	332.2	+ 0.3614		
UDD	334.0	- 0.4820		
KJN	334.8	- 0.0336		
UME	335.5	- 0.3136		
SKA	336.8	- 0.4511		
KIR	339.3	+ 0.2392		
Station	Distance (reduced to zero focus)	log <sub>lO</sub> (Amp) at 1 sec		
Station KJN	(reduced to zero	log <sub>10</sub> (Amp) at 1 sec		
	(reduced to zero focus)			
KJN	(reduced to zero focus) 97°1	- 0.0336		
KJN KIR	(reduced to zero focus) 97.1 98.7	- 0.0336 + 0.2392		
KJN KIR NUR	(reduced to zero focus) 97.1 98.7 99.4	- 0.0336 + 0.2392 - 0.1053		
kjn Kir Nur UME	(reduced to zero focus) 97.1 98.7 99.4 100.3	- 0.0336 + 0.2392 - 0.1053 - 0.3136		
KJN KIR NUR UME UPP	(reduced to zero focus)  97.1  98.7  99.4  100.3	- 0.0336 + 0.2392 - 0.1053 - 0.3136 + 0.3614		
KJN KIR NUR UME UPP SKA	(reduced to zero focus)  97.1  98.7  99.4  100.3  102.9  103.6	- 0.0336 + 0.2392 - 0.1053 - 0.3136 + 0.3614 - 0.4511		

Again for the earthquake on the 15 July, 1965, we have the following results at 1 cycle per sec frequency. (Obviously the data is not accurate to four decimal places!)

Station	Azimuth	log <sub>10</sub> (Amp) at 1 sec		
KLS	327 <b>°</b> 8	0.1578		
GOT	330.0	0.2462		
UPP	331.0	0.2488		
UME	334.4	0.1685		

SKA 335.4		- 0.4373		
KIR	338.3	0.7596		
Station	Distance	log <sub>10</sub> (Amp) at 1 sec		
KIR	90 <b>°</b> 1	0.7596		
UME	91.3	0.1635		
UPF	93.6	0.2488		
SKA	94.7	- 0.4373		
KLS	95•7	0.1578		
GOT	97.0	0.2462		

In the first example the back azimuths vary from 61° to 75° and in the second example from 67° to 74°. The change in back azimuths between the two examples is about 5.5° for each station.

The Fourier spectral program from seismogram to amplitude spectrum has been checked against an independent program. The spectral estimates seem therefore to be valid.

The problem remains - which of the assumptions we have made is not valid? Possible it is that the source function can vary very rapidly over very small azimuthal angles. In both the above examples the smallest and the largest amplitudes are next to each other in the distribution of azimuth. (This effect has not been checked on other data sets). If the source spectrum does vary so much with such small angles, then spectral analysis of short-period P-waves from earthquakes could only be done on a statistical basis. Explosions provide much more symmetrical sources, but their limited distribution prohibits their application to our present problem.

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The CDC 3600 computer of the Uppsala University Computer Centre was used for the calculations.

Table 1
Stations used (see also figure 1)

Station Code	Station Name	Location		
		Latitude	Longitude	
UPP	Uppsala	59°86n	17°63E	
UME	Une	63.82	20.24	
KLS	Karlskrona	56.17	15.59	
GOT	Göteborg	57.70	11.98	
UDD	Uddeholm	60 <b>.90</b>	13.61	
<b>6KA</b>	Skalstugan	63.58	12.28	
KIR	Kiruna	67.84	20.42	
KEV	Kevo	69.76	27.01	
SOD	Sodankylä	67.37	26.63	
NUR	Nurmijärvi	60.51	24.65	
KJN	Kajaani	64.10	27.70	

Table 2

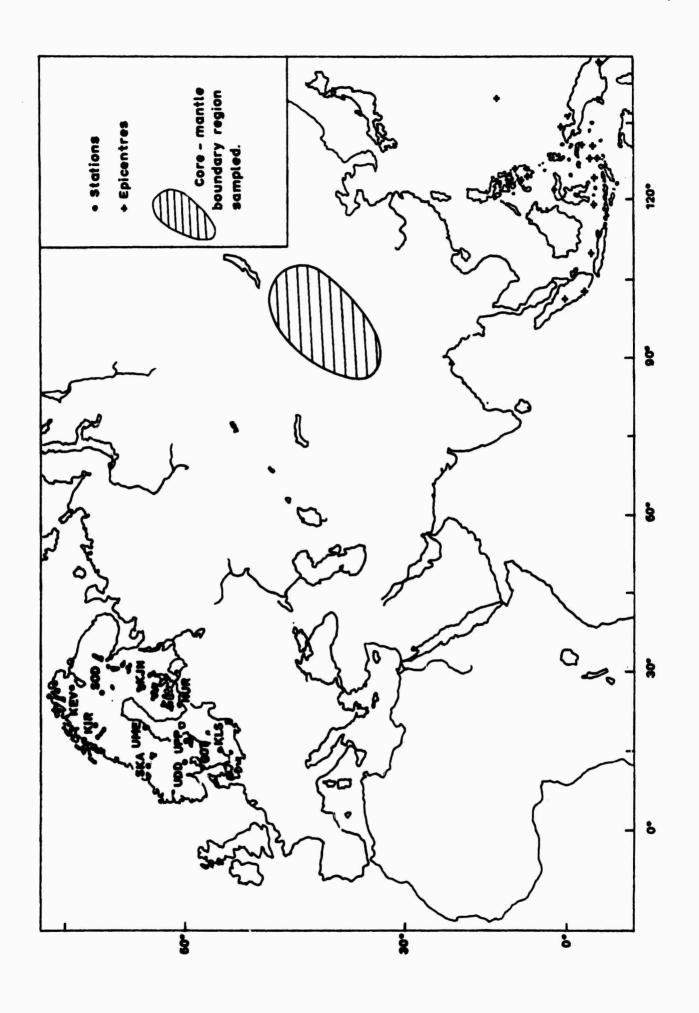
Earthquakes used (see also figure 1)

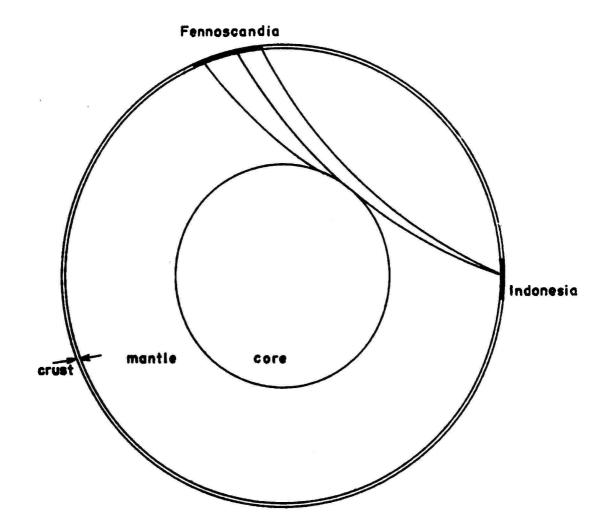
Data from USCGS

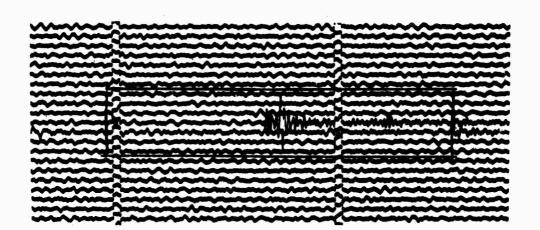
Date	Origin time	Origin time Epicentre		Depth	Magnitude
	(CMT)	Latitude	Longitude	(km)	(m)
	h m s				(UPP_KIR)
26.2.1963	20.14.08.7	<b>-</b> 7 <b>°</b> 5	146°2E	171	7.7
7.4.1963	22.36.03.4	-4.9	103.2	72	6.7
21.3.1964	03.42.19.6	-6.4	127.9	367	6.6
28.3.1964	11.30.09.8	0.5	122.3	140	6.2
8.7.1964	11.55.39.1	<b>-</b> 5•5	129.8	165	7.1
18.10.1964	12.32.24.1	-7.0	124.0	574	7.0
29.4.1965	15.48.57.1	<b>-5.</b> 6	110.2	504	6.3
15.7.1965	18.33.29.9	7•7	123.8	588	6.5
20.7.1965	13.18.27.4	7.5	124.3	45	5.9
20.8.1965	05.54.50.0	-5•7	128.6	326	6.7
21.8.1966	05.00.26.8	8.5	126.7	67	6.2
21.5.1967	18.45.11.7	-1.0	101.5	173	7.0
26.8.1967	00.36.42.1	12.2	140.7	33	6.6
24.5.1968	15.43.54.2	<b>-6.</b> 8	118.9	605	6.3
29.7.1968	23.52.15.0	-0.2	133.4	12	6.7
27.9.1968	03.58.55.1	-6.8	129.1	127	6.9

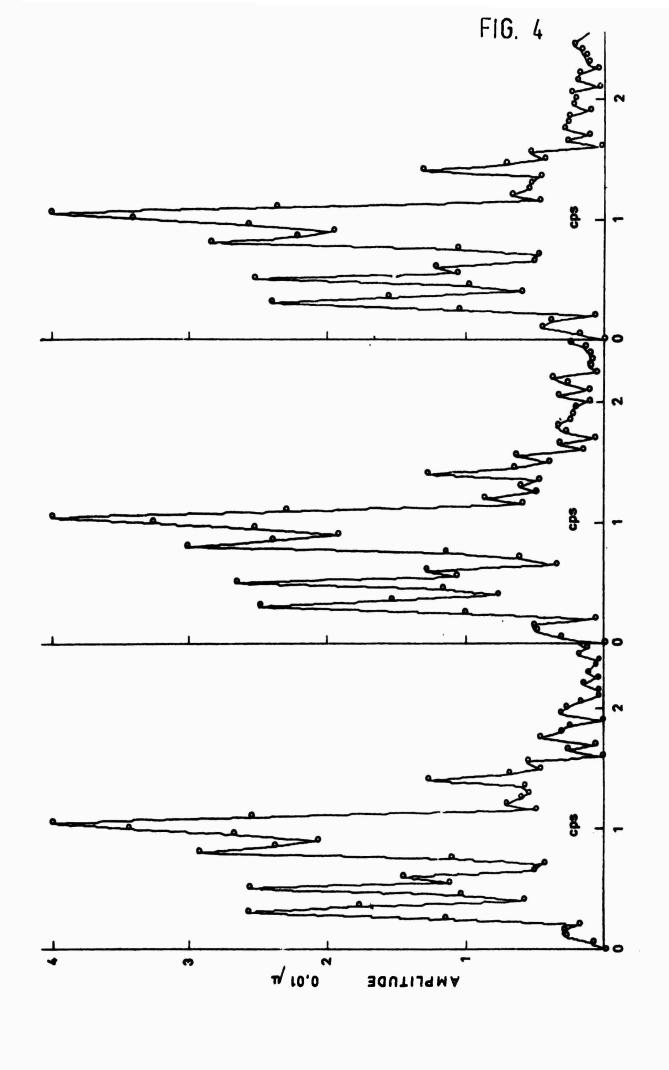
## Figure captions

- Fig. 1. Mercator projection of area of interest (showing Fennoscandian stations and Indonesian epicentres used).
- Fig. 2. Cross-section of the earth showing diagrammatically the ray paths from Indonesia to Fennoscandia (the diagram is merely suggestive and not to scale).
- Fig. 3. Short-period vertical-component P-wave recorded at Umea from the Banda Sea earthquake of 21 March, 1964.
- Fig. 4. The effect on the spectral amplitude of treating the record at Kiruna from the earthquake of 21 March, 1964, three times as a separate unit.
- Fig. 5. Plot of all raw data for 1 cps with station and epicentre terms removed (shows failure of method to give the expected result and also shows large scatter).

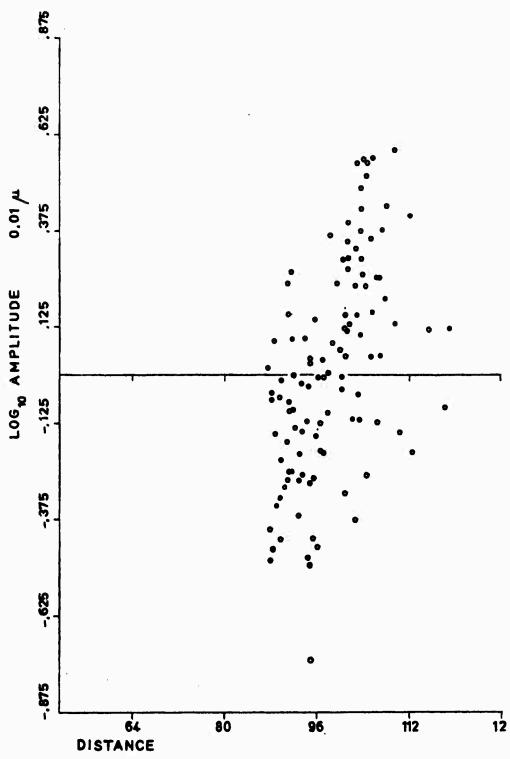












# Appendix

Program SOLVE is used to invert the system of linear equations described in the section "Analytical method". The program is not particularly efficient but is effective. Subroutine DIST 1 puts the earthquake-station pairs into their distance ranges. Subroutine SHUFFLE is used to vary the input without changing the coding of the original seismograms - earthquake 3 at station 6 has code 1316. Subroutine SELECT selects the frequencies required from the 20 frequencies of the input. Subroutine DIST 2 is here only formal but may be used to remove any prior estimate of station, distance or source terms.

```
MU08 . 10414C . ANSHLL . 3
*DEVAND.23000
AFOULD . TO ME
                               NOT REPRODUCIBLE
                                                                Page 1A
PETN. # . O.F.L.X
      PROGRAM SOLVE
      CALL WORK
      CALL EXIT
      SURPOUTINE WORK
      THIS PROGRAM COULD BE MORE EFFICIENT
ε
      MAXIMUM DIVENSIONS 110 RECORDS . 15 FREQUECTES
                                                               AND
1
      NST STATIONS . NPA PARAMÈTERS AND NEP EMICENTRES WHÉRE
       NST+NPA+NED IS LESS THAN OR EGUAL 10 49
      DIMENSION A(113,50).B(113,15).DISTANCE(110).FRQ(20).ANUM(60)
      DIMENSION X(50.15).AS(50.50).AT(125.20)
      DIMENSION INC(113) ATCOL(113) ATTITLE(6)
      FOUTVALENCE (AS. AT.)
      COMMON/10/44(50.50)
      READ 991 .NST.NEP.NPA.NRE.NER
  291 FORVAT(515)
      PRINT 992 .NST.NEP.NPA.NRE.NFR
  9#2 FORMATILX.15. + STATIONS. +.15. + EPICENTRES.
                                                       #, 15, # PARAMETERS.
         *.15.* RECORDS.
                           *.15,* FREQUENCIES. *)
      NCOL=NST+NEP+NPA+15 NROW=NRE+3
      READ 1. (FPC(I).I=1.NFR)
    1 FORMAT(15F5.2)
      KST=1+NST$KPA=1+NST+'4PA$JST=2SJPA=2+NST&JEP=2+NST+NPA
      KEPHNCOL
      ITITLE(1)=84FREO.
      DO 505.1=3.6
  505 ITITLE(I)=8H
      0.0= (L.!)4 TUG
C
      DO 2, 1=1, NROW
      DO 3.J=1.NCCL
      A(I,.J)=0.0
    3 CONTINUE
    2 CONTINUE
      PUT IN VALUES OF FIRST COLUMN O A(I, J) AND BOTTOM THREE ROWS OF
C
      A AND B
2
      DA 4.1=1.NOF
    4 A(1,1)=1.0
      NRFT= NRF+1
      DO 5.1=JST.KST
    5 A(NRFT.1)=1.0
      NRET= VRET+2
      DO 6.1=JEP.NCOL
    6 A(NRFT.1)=1.0
      NRFT=NRET-1
      DO 9991, I=JPA, KPA
 9991 A(NRET, I)=1.0
```

NRFT=NRF+1 DO 7.J=1.NFR A CONTINUE

7 CONTINUE ANUM(I) GIVES THE NUMBER OF NON-ZERO ELEMENTS IN EACH COLUMN OF A PPINT 12

12 FORMAT(4X, +INPUT DATA+)

PRINT 14

14 FORWATI#

PRINT 15. (FRG(I).I=1.NFR)

FORMAT(9x. #FREQUENCY #.15F7.2)

PRINT 15

16 FORMATCIX. # EPI STAT DISTANCE\*)

DO 9. I = 2. NCOL

9 ANUMITIED.D

ANUM(1)=FLOATF(NRE)

READ IN DATA C

DO 100, I=1, NRF

READ 10.MEPIC.MSTAT.DISTA 10 FORMAT(1X,212,F6.1)

DISTANCE(I)=DISTA

MEPIC=MEPIC-10 SMSTAT=MSTAT-10

MOUMPENSTAT

CALL SHUFFLF (MEPIC, NEP, MSTAT, NST)

MEMSTATHE SNEMEDIC+KPA

A(1, M)=1.0

A(1,N)=1.0

ANUM(4)=ANUM(4)+1.0

ANUW(N)=ANUW(N1+1.0

CALL DISTI(DISTA, NPA, KDIST)

KA=KDTST+1+NST

IDR(I)=KA

A(1,KA)=1.0

ANUM(KA)=ANIJM(KA)+1.G

PFAD 23, (FRQ(K), K=1,20)

23 FORVAT(5F12.4)

CALL SFLECT (FRO.NER) FRO IS HERF THE INPUT AMPLITUDE DATA τ

PRINT 24, MEPIC, MSTAT, DISTA, (FRQ(K), K=1, NFR)

24 FORWAT(1X,215,F8.1,11F10.6)

Z=DISTANCF(I)

SUBTRACT FITTED CURVE, SPECIFIED IN DIST2, AND INST EFFECT C

CALL DIST2(FRO.Z.NFR.MSTAT.NST)

DO 9981.K=1.NFR

QUAL B(I,K)=FRQ(K)

IND CONTINUE

HAVE NOW FED DATA INTO B AND PARAMETERS INTO A. AND SUBTRACTED

FITTED CURVE AND INST EFFECTS FROMB

```
LUUT THIEL
INAL FORMAT (# VATRIX A*)
      00 1003.T=1.NPCW
      DRINT 1002, (4(1,J), J=1, NCOL)
                                                                Page 3/2
 INDA CONTINUE
 inna FORMAT(1x,50F2.0)
      PRINT 14
      PRINT 1004
 1004 FORMATE #
                  MATPIX 8#1
      DO 1005 + T=1 + NROW
      DPINT 1006, (3(1,J), J=1,NFP)
 INGS CONTINUE
      FORMAT (1X+15F8-3)
 1006
      DRINT 14
        WE NOW FORM ATA AND ATE SATA IS INAA AND ATE IS IN X
C
      20 10 4+ T=1+NCOL
      no 105.J =1.1
      AA(T. J)=0.0
      20 105.K=1.NROW
  1 6 00(1.J)=00(1.J)+4(K.1)+A(K.J)
      (L+T) AA=(L+T)ZA
  105 CONTINUE
      1 = 1 - 1
      nn 107. J=1.L
      ALIJ.T)=AA(T.J)
  107 AC(J.T)=AA(J.T)
  194 CONTINUE
      00 110 +T=1+NCOL
      nolil . J=1 . NER
       N.C=(L,1)X
      DO 112+K=1.490W
  112 X(I,J)=X(I,J)+A(K,I)*B(K,J)
  111 CONTINUE
  JIN CONTINUE
      DOINT 1010
 INTO FORMAT(# MATRIX ATA#)
      DO 1011.I=1.NCOL
      DRINT 1002+(44(1+J)+J=1+NCOL)
 INTE CONTINUE
      DETERM=0.0
      NMAX=50
      CALL MATINY(NCOL, X, NFR, DETERM, NMAX)
      DRINT 9992, DETFRM
 0002 FORMAT (1X. *DETERM = +. E12.4)
      62F=0.0
```

NOT REPRODUCIBLE

DO 2001+1=1+NCOL DO 2002+J=1+NCOL

VADEO.0

```
DO 2003,K=1,NCOL
 ?n 3 VAR=WAR+AA(T,K)+AS(K,J)
      TF(T.EQ.J) VAR=VAR-1.0
                                     NOT REPRODUCIBLE
                                                                    Tage 44
      VAR=VAR*VAR
 2002 IF(VAR.GT.GRE) GRE=VAR
 2001 CONTINUE
      DRINT 14
      CRESCRIF (CRE)
      DRINT 2005,GRF
20 5 FORMAT( * GREATEST ERROR IN PRODUCT OF MATRIX AND ITS INVERSE = *.
     1512.4/1
      DO 120.1=1.NROW
      OO 121 ,J=1,NFR
      DIMED.O
      DO 122. K=1.NCOL
  122 DUM = DUM + A(I,K) + X(K,J)
      AT(I,J)=DUM-B(I,J)
      (L,I)TA*(L,I)TA=(L,I)B
      AT(I \bullet J) = X(I)R(I) \bullet J) - AT(I \bullet J)
  121 CONTINUE
  120 CONTINUE
      DO 500.J=1.NER
      DO502, I=1,NRF
  502 ATCOL(I)=AT(I,J)
      TTJTL = (2)=8H
      CALL ENCODE (ITITLE)
      CALL FMTS(8)
      CALL FMTI(J,1)
CALL GRAPHI(DISTANCE, ATCOL, -NRE, 3H7X8, 4HAUTO, ITITLE, LOHDISTANCE = 7,
     15HAMP..,5260606060606060B)
  500 CONTINUE
    X CONTAINS THE SCLUTIONS AND B CONTAINS THE ERROR SQUARED
      WE FIND NOW THE STANDARD DEVIATIONS
C
      SQAN1 = SQRTE (ANUM(1)-1.0)
      SOAN2 = SQRTF (ANUM(1) - NPA+1.0-1.0)
      ANNN=ANUM(1)-1.0+(NPA+NST+NEP+1)
      ANNN =MAX1F(ANNN+1.0)
      SOANS = SOPTF (ANNN)
      nn 150,1=1,NFR
      AA(I,1)=0.0
      DO 151 .J=1.NRF
  1 \le 1 \le AA(1,1) = AA(1,1) + B(J,1)
      DUM=SORTF(AA(I+1))
      AA(T,1)=DUM/SQAN1
      AA(I, 2)=DUM/SQAN2
      AA(1,3)=DUM/SOAN3
  150 CONTINUE
      EXCEPT FOR FIRST COL ANUM(I) IS NO IN COL MINUS ONE
       no 159 J=1,NCOL
      K=J+3
```

The second section of the sect

L=J+1

no 160, I=1, NFR

```
AA(I,K)=0.0

DO 161, M=1.NPE

161 AA(I,K)=AA(I,K)+B(M,I)#A(M,L)

160 AA(I,K)=SGRTF(AA(I,K)/ANUM(L))
```

Page 5A

```
. FO CONTINUE
    PRINT 14
PRINT 200
2 0 FORMAT(#
               RESULTS#/)
    TOS THIPS
201 FORMAT(* CONSTANT*)
202 FORMAT(1X.16F8.4)
    nn 204 ,J=1,3
    PPINT 202.(X(1,1),AA(1,J),I=1,NFR)
204 CONTINUE
    DRINT 14
    ODINT 205
205 FORMAT(1X. #STATIONS#)
207 FORMAT(1X.12)
    no 206 .I=2.KST
    J=1-1 $L=1+2
    DPINT 207.J
    PRINT 202, (X(I,K),AA(K,L),K=1,NFR)
206 CONTINUE
    DRINT 14
    PRINT 208
208 FORMAT(1X. *PARAVETERS*)
    DO 200 ,1=JDA,KDA
     J=1-1-NST 4 L=1+2
    DOTNITZOT, J
    PRINT 202. (X(I.K).AA(K.L).K=1.NFR)
209 CONTINUE
     PRINT 14
     PRINT 211
211 FORWAT(1X, *FPICFNTRES*)
    DO 212 ,I=JFP,KEP
    J=1-1-NST-NPA SL=1+2
    DRINT207.J
    PRINT 202.(X(I,K),AA(K,L),K=1,NFR)
212 CONTINUE
    PETHON
    END
    SUBROUTINF DISTI(DISTA, NPA, KDIST)
   X IS DISTANCE TO RIGHT OF FIRST INTERVAL, N IS NUMBER OF FIRST LONG
   1 INTERVAL, Y IS THE DISTANCE BETWEEN THE TOP OF FIRST AND BOTTOM
   1 OF FIRST LONG INTERVAL, W IS WIDTH OF SHORT INTERVAL , LONG IS 5.
    x=101.0 $N=4 $Y=4.0 $W=2.0
    Z=DISTA-X
    IF(Z.LT.0.0)Z=0.125
```

```
IF(Z.GT.Y) GO TO 1
      7=7/W
      KDIST=INTF(7)+1
                                                                     Page 64
      PFTUPN
    1 Z=(Z-Y)/5.0
      KDIST=INTF(7)+N
      RETHEN
      END
      SUBROUTINF DIST2(FRQ,Z,NFR,MSTAT,NST)
      DIMENSION FRO(1)
      GO TO 2
    2 CONTINUE
      RETHRN
      END
      SUBROUTINE SHUFFLE (MEPIC, NEP, MSTAT, NST)
      DIMFAGION NUTT(16), NURT(11)
      SORTS EPICENTRES AND STATIONS INTO NUMBERED ORDER
      DATA((NUTT(1), I=1,16)=7,C,O,O,1,C,C,O,0,0,2,3,0,5,0,4),((NURT(1),I=1,
     1,111=1,2,3,4,5,6,7,8,9,0,10)
      WEDIC=NUTT(WEDIC)
      MSTAT=NURT(WSTAT)
      RFTURN
      FND
      SUBROUTINE SELECT (FRO, NFR)
      DIMENSION FRO(1).FL(20)
      SELECTS THE FREQUENCIES REQUIRED FROM THE 20 READ
      DATA((NUM(I), I=1,10)=1,2,3,4,5,6,7,8,9,10)
      DO 1 I=1.20
    1 FL(!)=FQQ(!)
      DO 2 1=1.NER
    2 FPC(I)=FL(NUM(I))
      PFTIPN
      FND
         SCOPE
PLOAD,
( 1, 12()+)>
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